

A neural network based method for input parameter selection



Stefan Lotz^{1,2}, Jacques Beukes^{2,3}, Marelle Davel^{2,3}

¹SANSa Space Science Directorate, Hermanus, South Africa

²Multilingual Speech Technologies (MuST), North-West University, South Africa

³Centre for Artificial Intelligence Research (CAIR), South Africa



Introduction

- NNs yield predictions, without aiding understanding of input-output relationship
- Fully connected networks mix signal from all inputs as information flows through the network
- Input parameter selection usually done by the user, outside NN training framework

- Can we configure a NN to allow for separation of inputs in to subsets?
- Can we use this to find a ranking of input parameters in terms of importance?

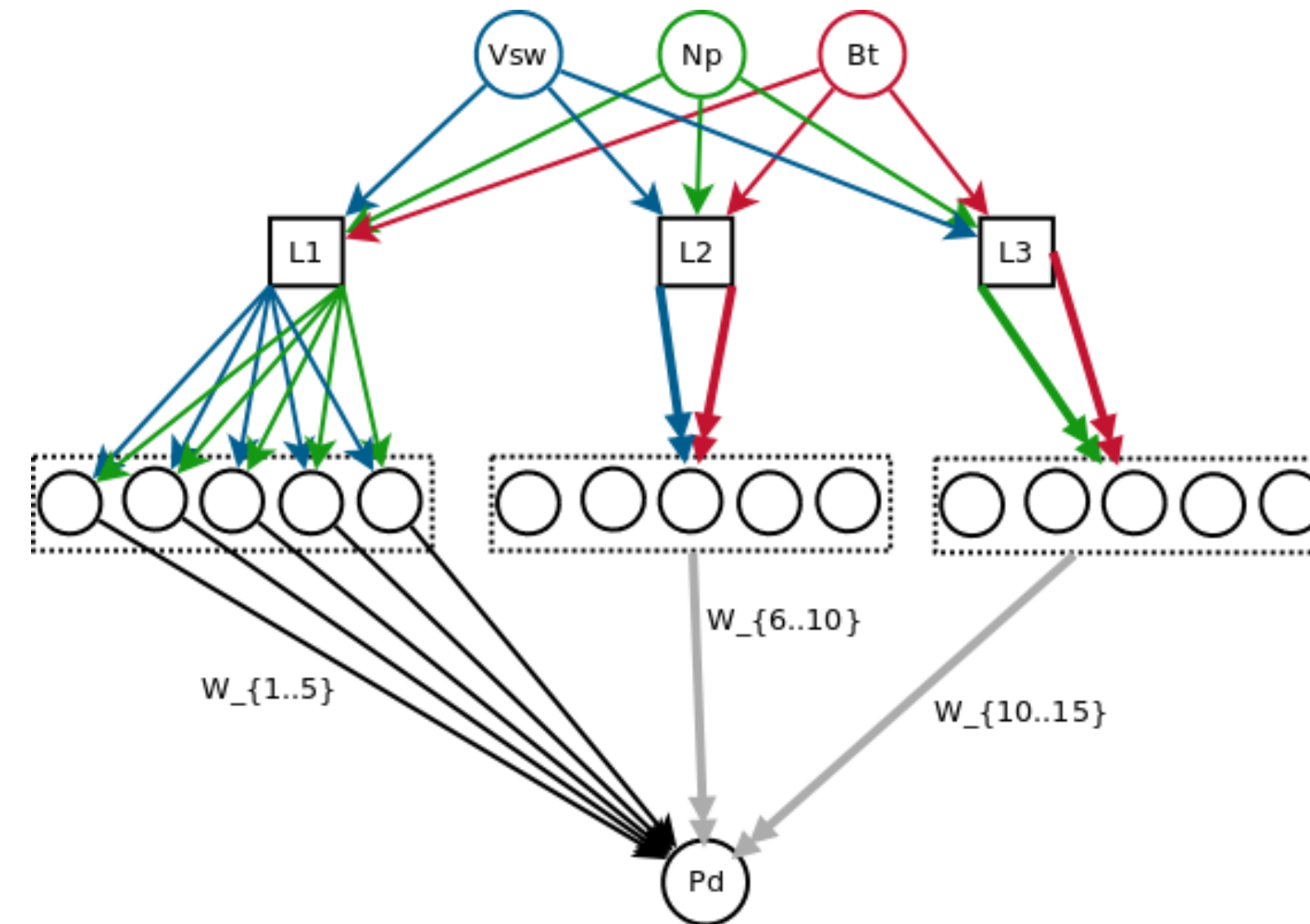
→ We present a first try: pair-wise inputs through λ -layers

A pair-wise input NN

Toy problem: Predict

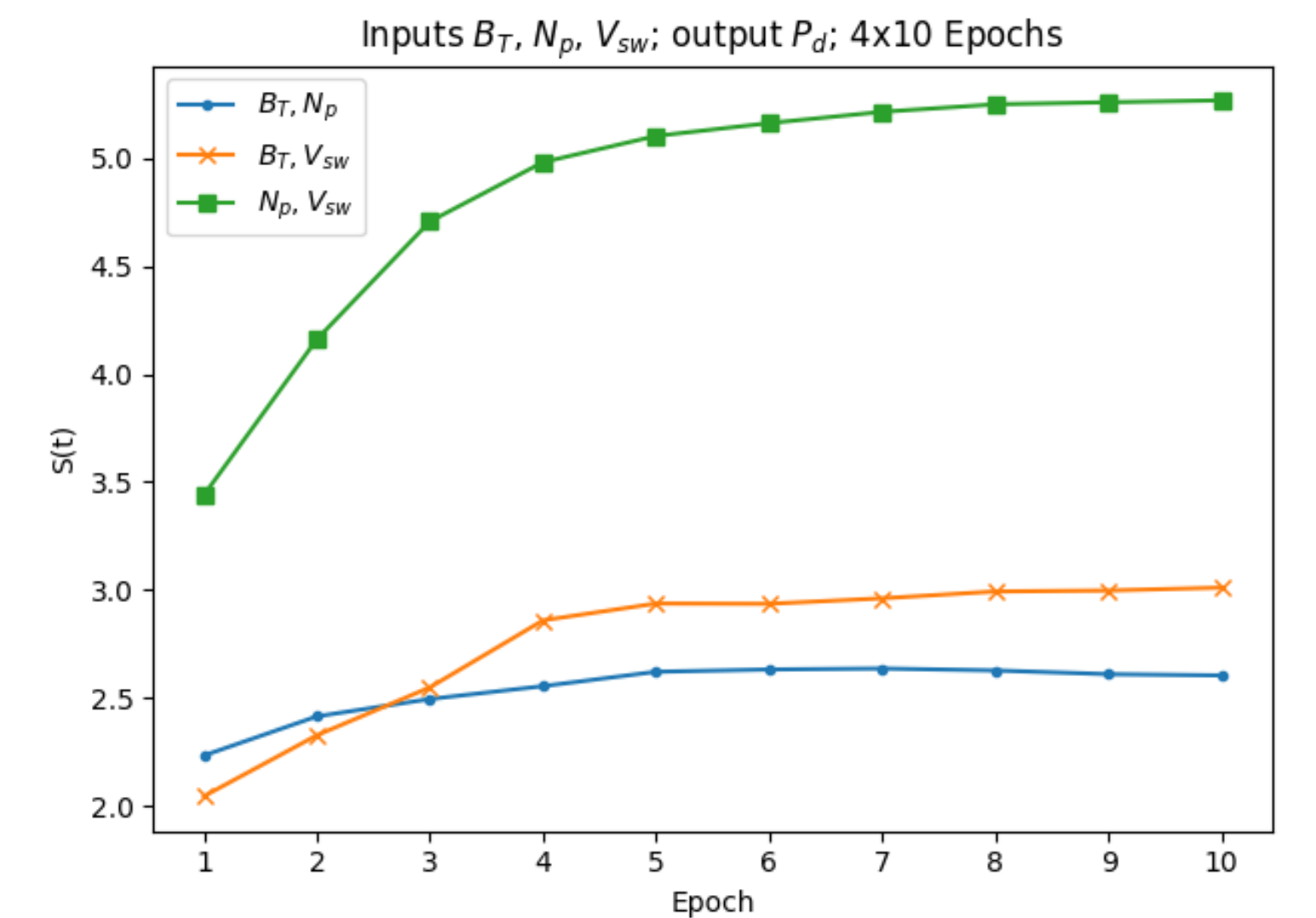
$$P_d \sim N_p V_{sw}^2$$

from V_{sw} , N_p , total IMF B_T



Track sum of normalised weights $W_i^*(t)$ at every training epoch for the pairs of inputs

$[V_{sw}, N_p]$ dominates as expected



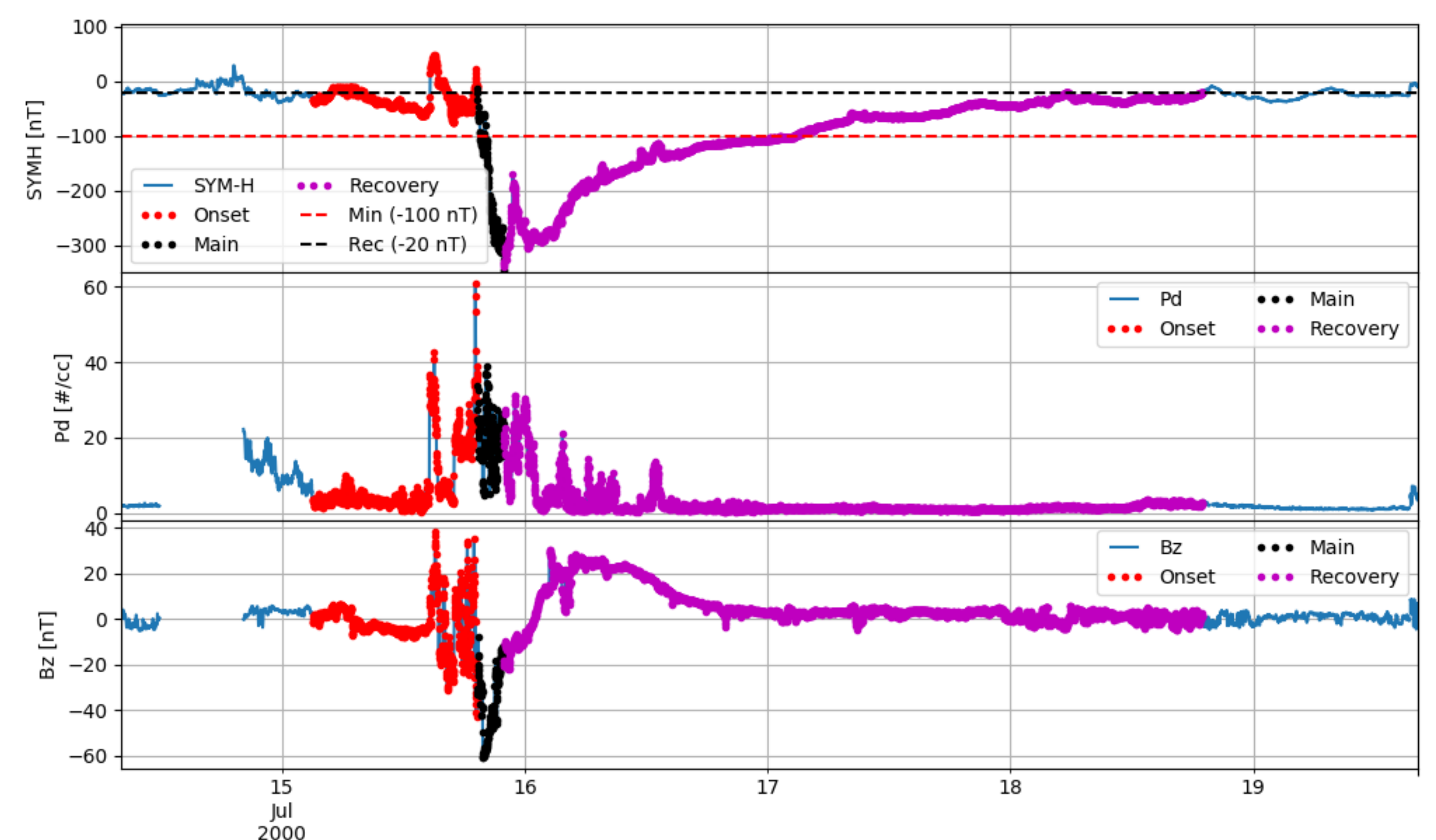
Predict SYM-H with storm phase information

- Dst / SYM-H prediction from solar wind input has been fairly successful [e.g. 1]
 - Storm phase information could be important source of information during training
- We develop simple FFNN model to predict SYM-H from solar wind parameters, with and without phase information

Data Set

- Interval 2000 – 2018, Inputs: OMNI 1-min, Output: SYM-H
- SYM-H < -100nT must be crossed, recovery at -20nT
- 97 storms identified, $N = 396,164$ minutes of data (error-free)
 - Training (TRN): 67 Storms, $N = 282,517$ (71.3%)
 - Validation (VAL): 15 Storms, $N = 57,634$ (14.1%)
 - Out of sample test (TST): 15 Storms, $N = 56,013$ (14.5%)
- No mixture of events → Independent TRN/VAL/TST sets
- Storm phases encoded with 100 - Onset | 010 - Main | 001 - Recovery

Storm and Phase identification in SYM-H

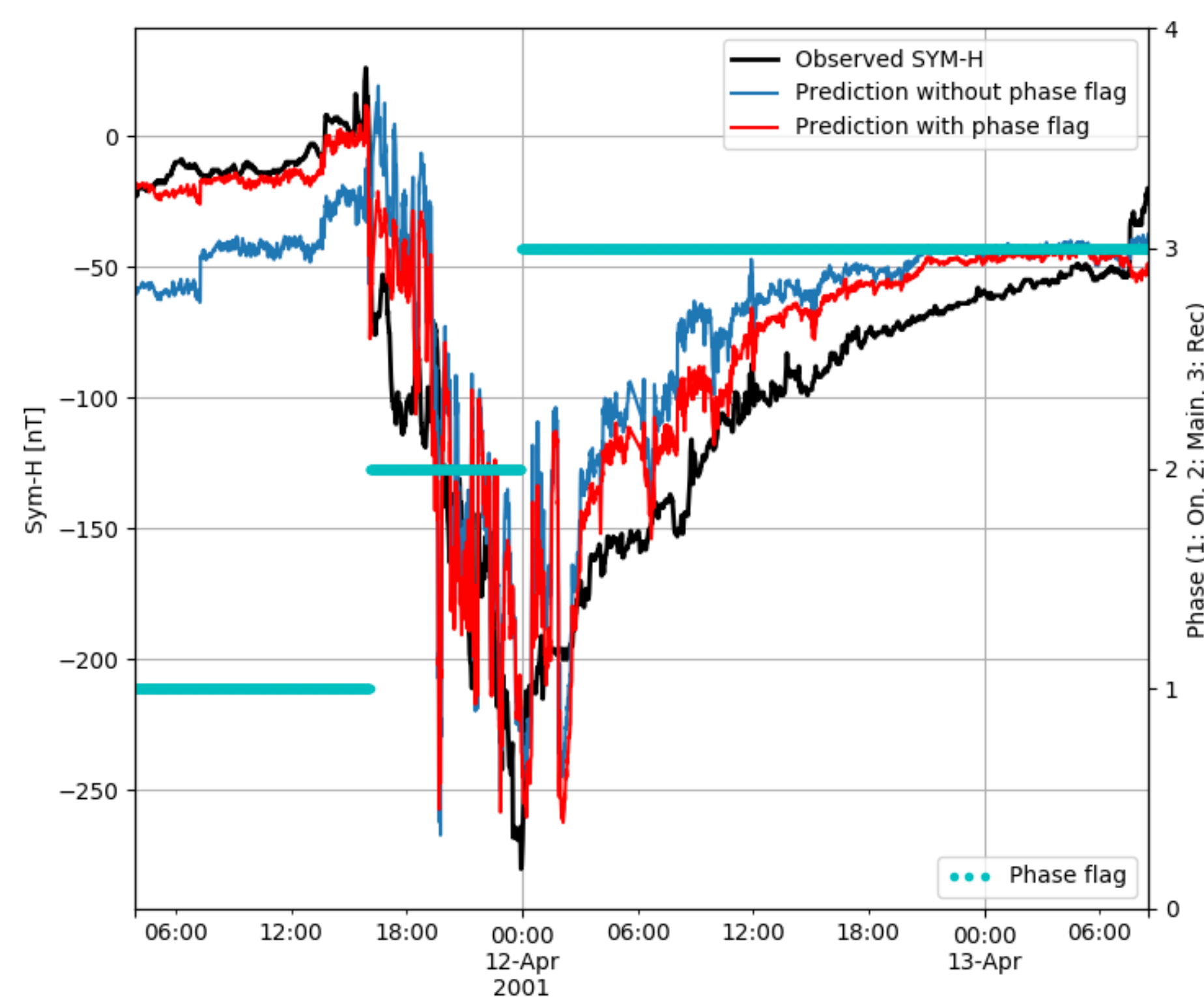


Geomagnetic storm intervals selected from SYM-H [See 2].

FFNN Model

- No phase**
- Inputs (at t and $t - 180$) V_{sw} , N_p , P_d , E_m , B_T , B_x , B_y , B_z
 - 16:50:1 FFNN
 - relu activations
 - adam optimiser
 - batch_size = 64
 - Performance (R) on TST 0.78

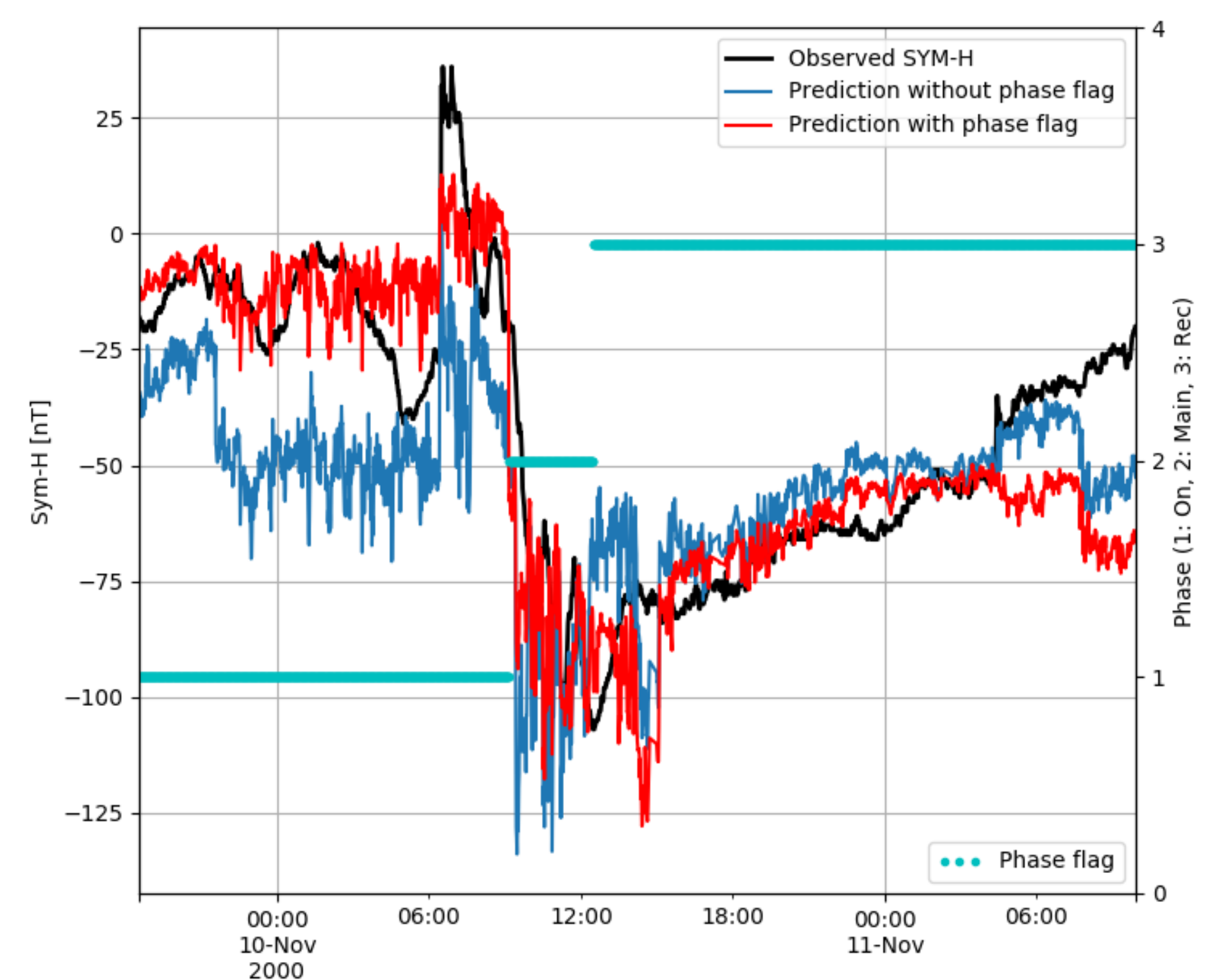
No phase



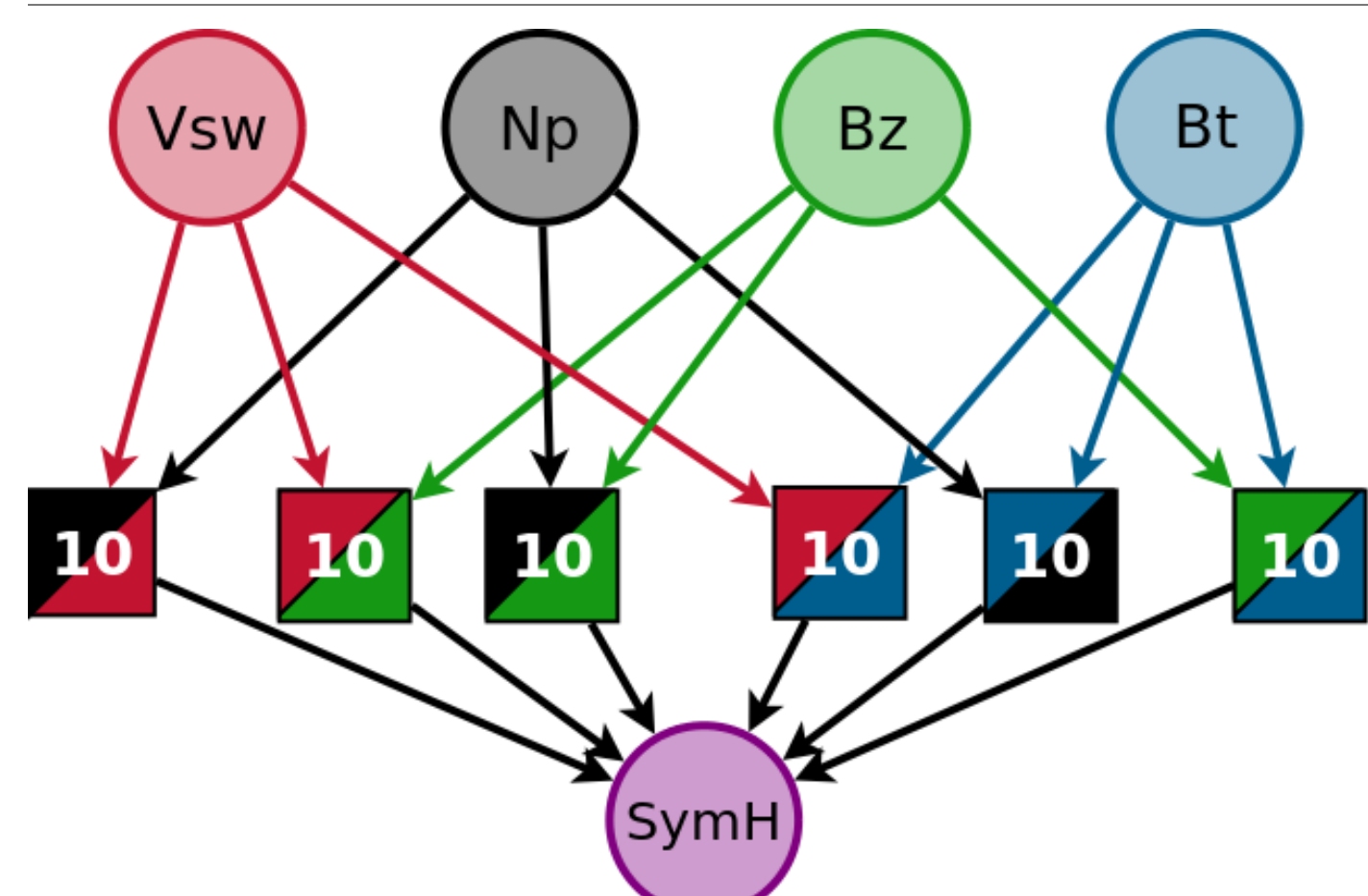
FFNN Model

- With phase**
- Inputs (at t and $t - 180$) V_{sw} , N_p , P_d , E_m , B_T , B_x , B_y , B_z
 - Phase one-hot encoding: 100-Onset | 010-Main | 001-Recovery
 - 22:50:1 FFNN
 - relu activations
 - adam optimiser
 - batch_size = 64
 - Performance (R) on TST 0.85

With phase

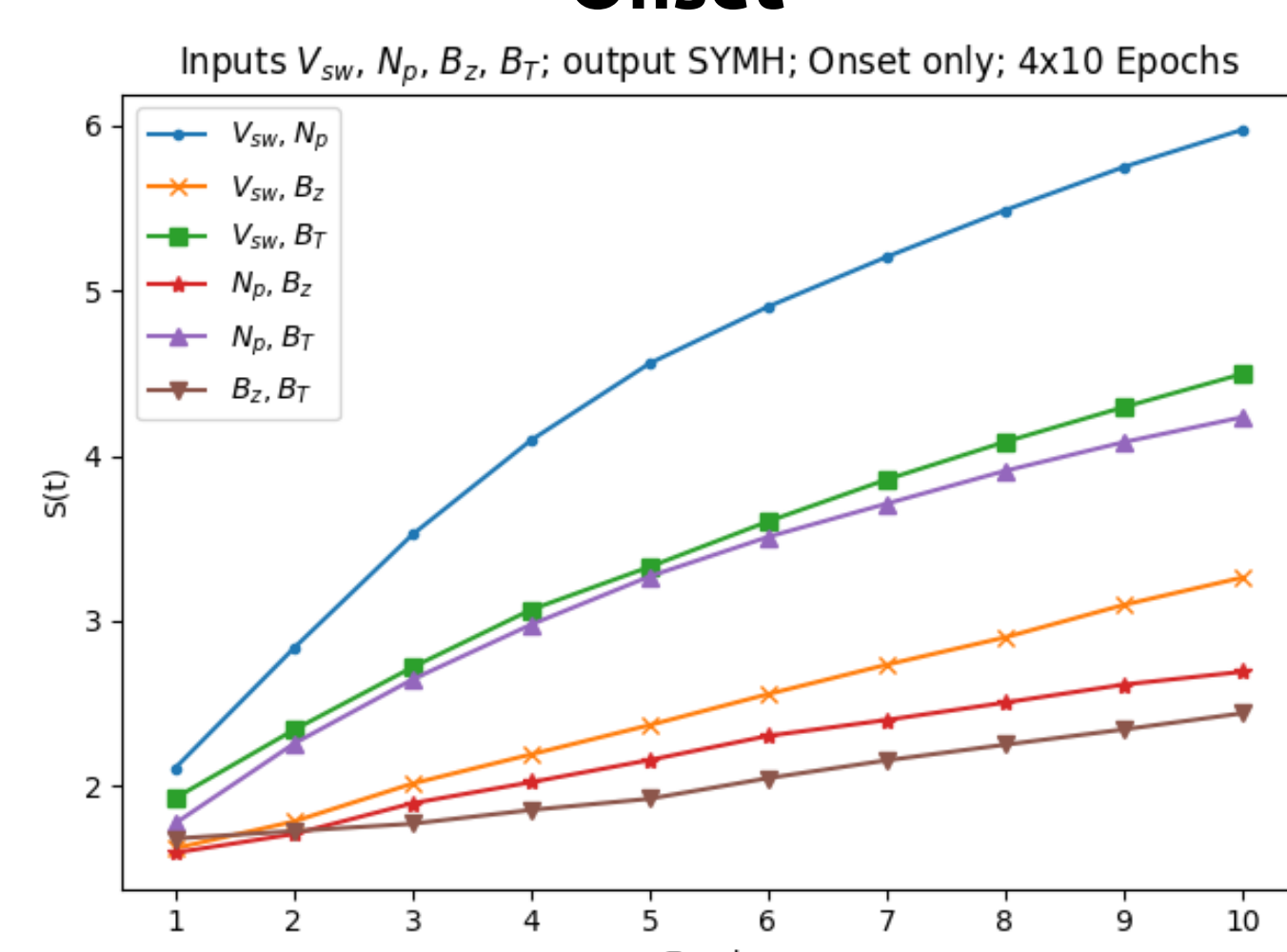


Parameter Selection per Storm Phase



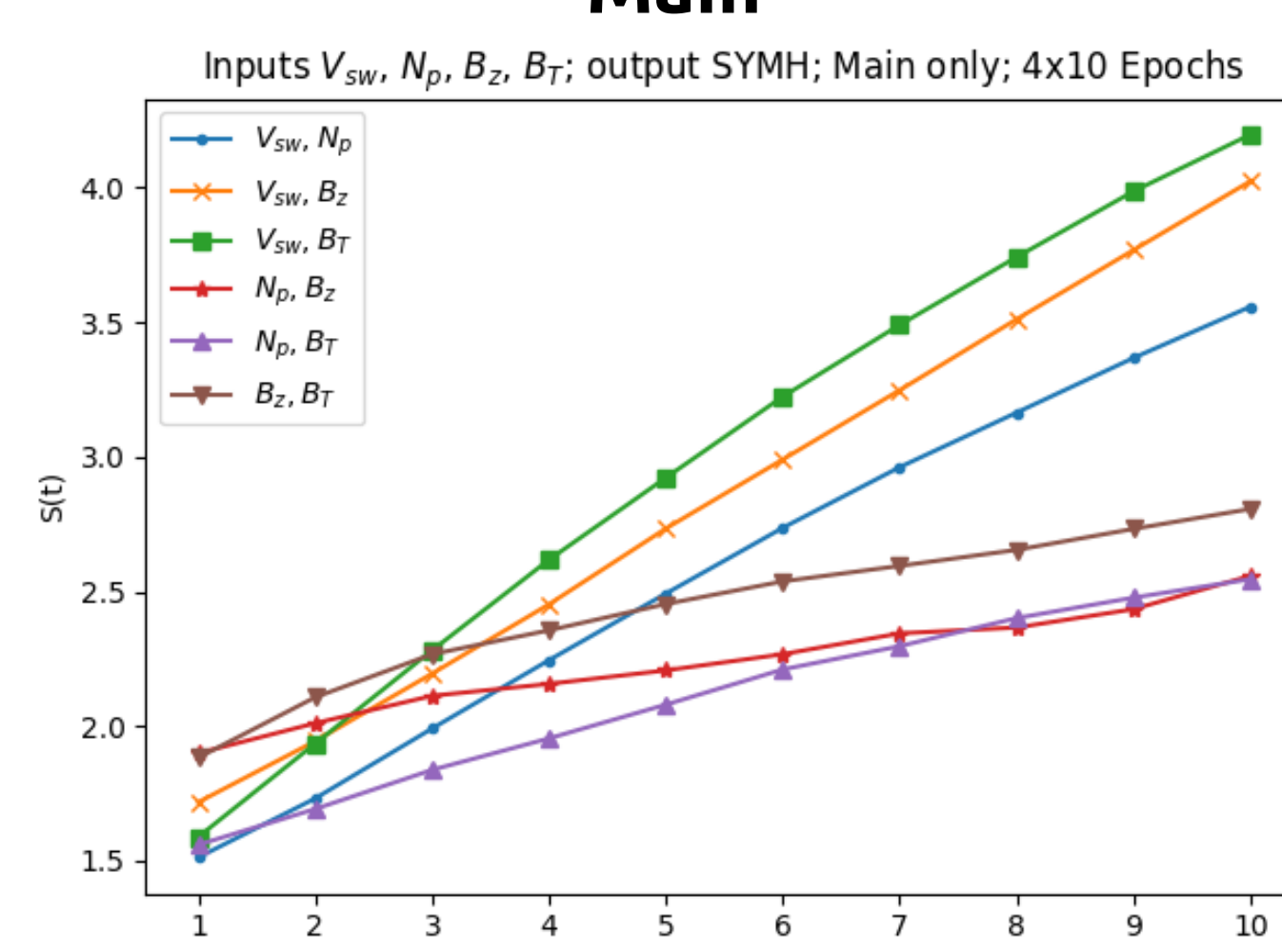
- 4:60:1 FFNN with pairwise λ -configuration
- Use reverse rank to score each input
- Conclusion: V_{sw} is always influential, N_p not important during main phase, but IMF B_T , B_z is

Onset



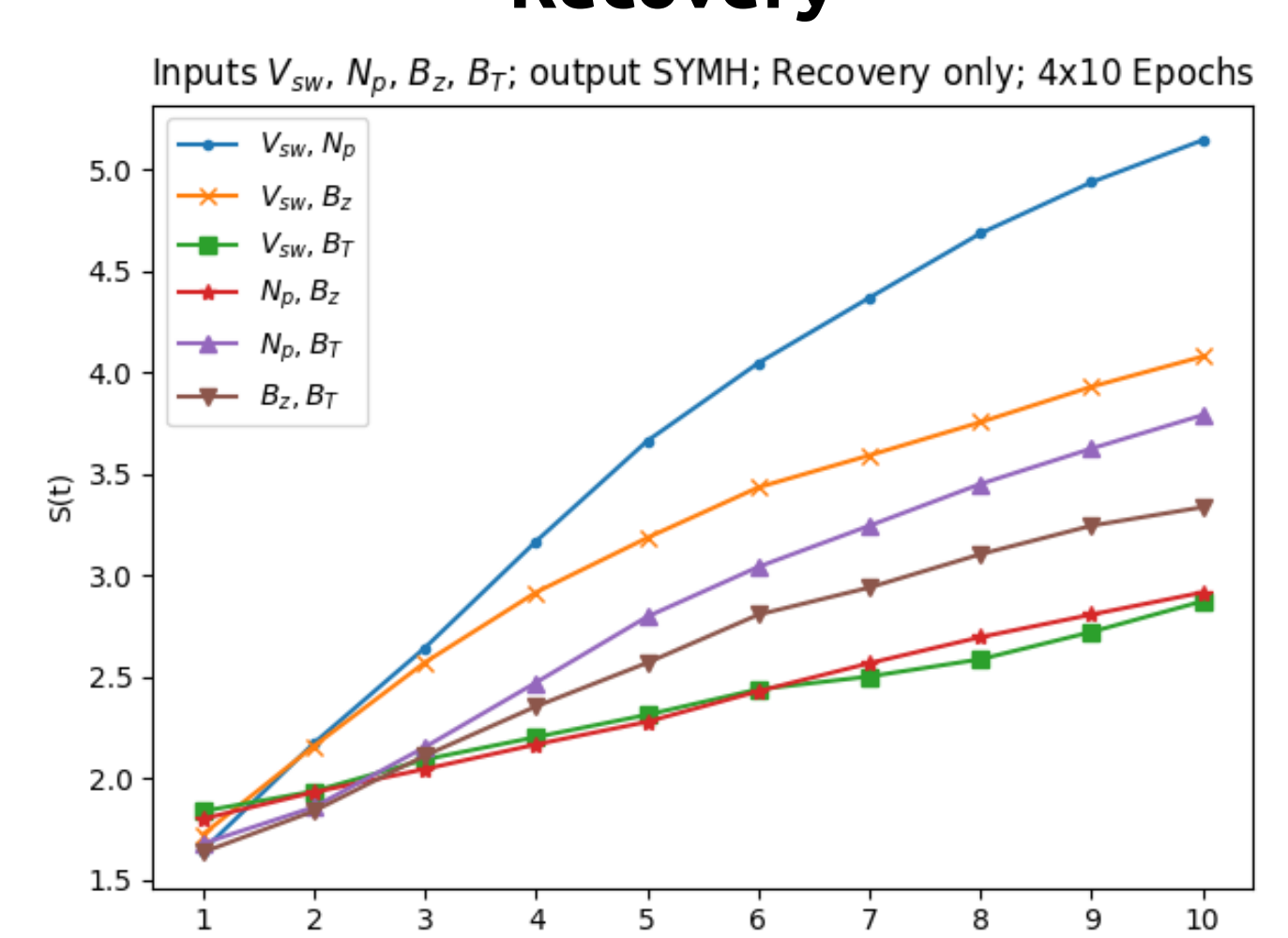
#	Par	Ranks	Scores	Tot Score
1.	V_{sw}	1,2,4	6,5,3	14
2.	N_p	1,3,5	6,4,2	12
3.	B_T	2,3,6	5,4,1	10
4.	B_z	4,5,6	3,2,1	6

Main



#	Par	Ranks	Scores	Tot Score
1.	V_{sw}	1,2,3	6,5,4	15
2.	B_T	1,4,6	6,3,1	10
3.	B_z	2,4,5	5,3,2	10
4.	N_p	3,5,6	4,2,1	7

Recovery



#	Par	Ranks	Scores	Tot Score
1.	V_{sw}	1,2,6	6,5,1	12
2.	N_p	1,3,5	6,4,2	12
3.	B_z	2,4,5	5,3,2	10
4.	B_T	3,4,6	4,3,1	8

References

- [1] M. A. Gruet, M. Chandorkar, A. Sicard, E. Camporeale. *Multiple hours ahead forecast of the Dst index using a combination of Long Short-Term Memory neural network and Gaussian Process*. Space Weather (2018), doi: 10.1029/2018SW001898.
- [2] S. I. Lotz and D. W. Danskin. *Extreme value analysis of induced geoelectric field in South Africa*. In: Space Weather 15 (2017), doi: 10.1002/2017SW001662.